The background of the right half of the slide is a photograph of Earth from space. The sun is visible in the upper right corner, creating a bright lens flare and illuminating the Earth's surface. The horizon of the planet is visible, showing the curvature of the Earth and the blue of the atmosphere.

The Electrode-less Z-Pinch as a metrology source in the 20-50 nm range.

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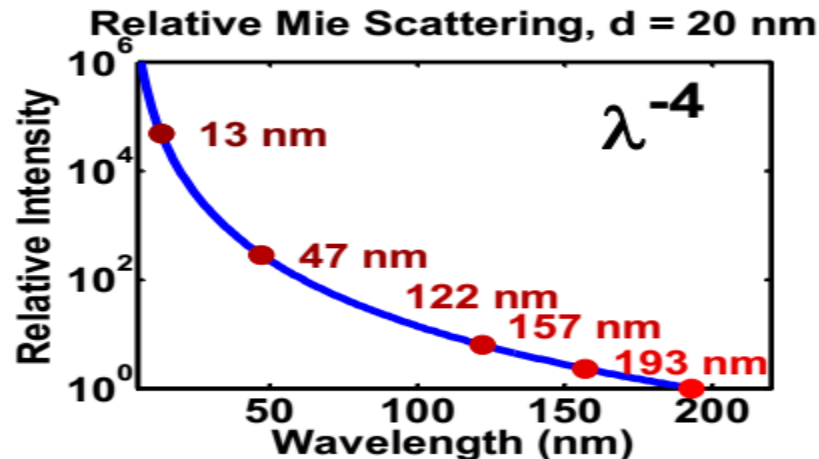
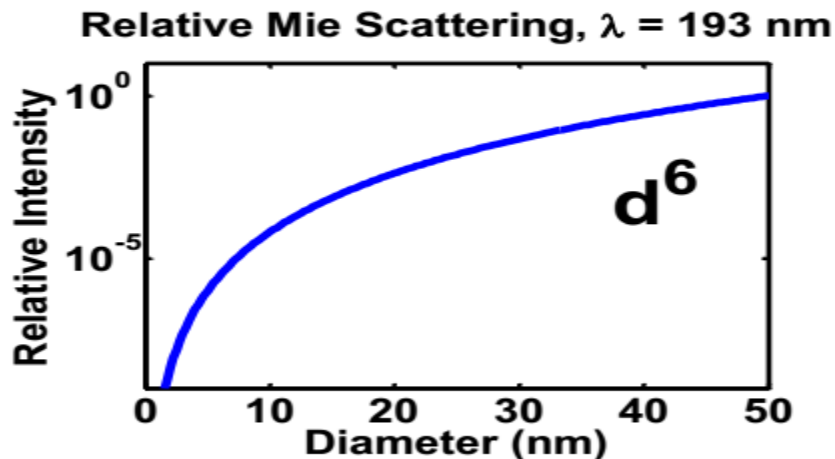
Outline

- Why these wavelengths?
- Theoretical approach
 - How? Brief review of EQ-10
 - Possible gasses
 - Energy balance
- Reality
 - Diagnostics
 - Results
 - Interpretation
- Conclusions

Why does anybody care about 20-50 nm light?

- Most patterned wafer inspection systems rely on broad-band light (down to ~ 200 nm) or lasers (ArF, 193 nm) to detect defects.
- As critical dimensions have been shrinking, so has the size of a so-called “killer defect” that must be detected. Defects are sub-wavelength, so can’t be effectively imaged.
- While the detailed interaction of the illuminating light with the defect is arbitrarily complicated, one can consider Mie scattering as a simpler, perhaps relevant model

Mie scattering amplitude from a particle $\sim d^6/\lambda^4$ (d particle size, λ wavelength)



Trends in relative scattered intensity of a sphere in free space as functions of diameter, d , and wavelength, λ .

About 200X more signal at 47 compared to 193 nm

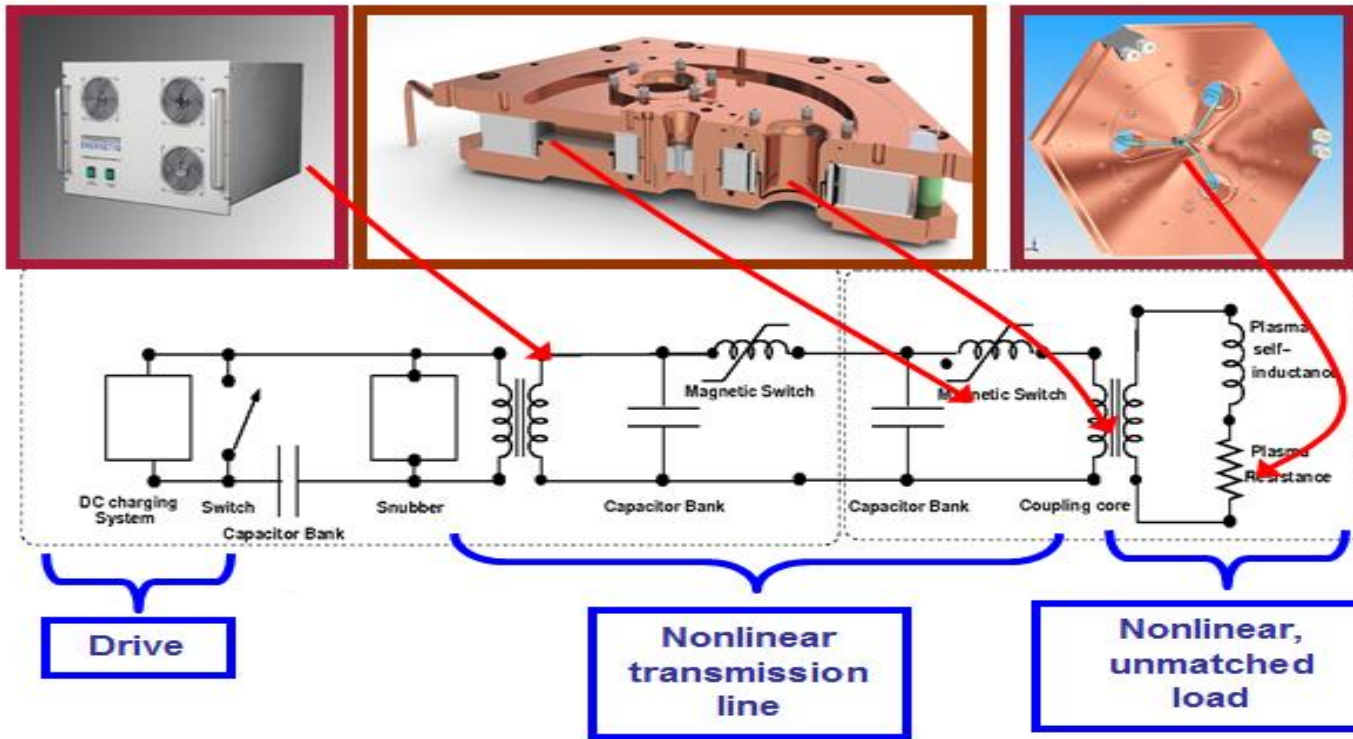
Barnes, Bryan M., et al. "Assessing the wavelength extensibility of optical patterned defect inspection." *SPIE Advanced Lithography 2017: Metrology, Inspection, and Process Control for Microlithography XXXI*. Vol. 10145. 2017.

How do we make the light?

If all you have is a hammer, everything looks like a nail..

If all you have is a Z-pinch, everything looks like a plasma...

Simplified schematic diagram of power system and source.



Horne, Stephen F., et al.
 "Development of a High Pulse Rate EUV Source." *Proceedings of SPIE, the International Society for Optical Engineering*. Society of Photo-Optical Instrumentation Engineers, 2009.

Available from
www.energetiq.com

Z-pinch in Xenon - a baseline

- Source produces $\sim 20\text{W}$ inband ($13.5\text{ nm} \pm 1\%$, 2 pi) in Xenon; $\sim 60\text{ W}$ at 11 nm
- $T_e \sim 20\text{ eV}$
- Plasma starts out at $\sim 100\text{ mT}$ gas, 1 cm long by 3 mm radius cylinder; $N \sim 3.5e21/\text{M}^3$; about 10^{15} atoms in the pinch
- Compresses to $\sim 0.2\text{ mm}$ radius; requires $\sim 1.4\text{ J/pulse}$; density increase of ~ 200
- Energy/radiating particle - $\sim 1.4\text{ J}/10^{15}\text{ particles/q} \approx 9\text{ keV/particle}$
- Energy required to ionize Xenon from neutral all the way to $\text{Xe}^{+10} \sim 800\text{ eV}$
- Ratio – takes about $10\times$ theoretical minimum to produce the radiating particle.

If we want to make 20-50 nm, we need to choose a working gas

- NIST spectral database used;
http://physics.nist.gov/PhysRefData/ASD/lines_form.html
- Search for lines with high relative intensity. Neon, Argon look like candidates... NIST tables more complete for Neon...
- Ideal radiative state will have an ionization barrier – works for 2.88 nm in Nitrogen for Water Window microscopy (Helium-like nitrogen) but nature isn't always so kind...
- Neon has many lines in this range – from 20 nm to 50 nm.
- Ne +3 (IV) looks promising...

NIST will estimate a spectrum – but beware...

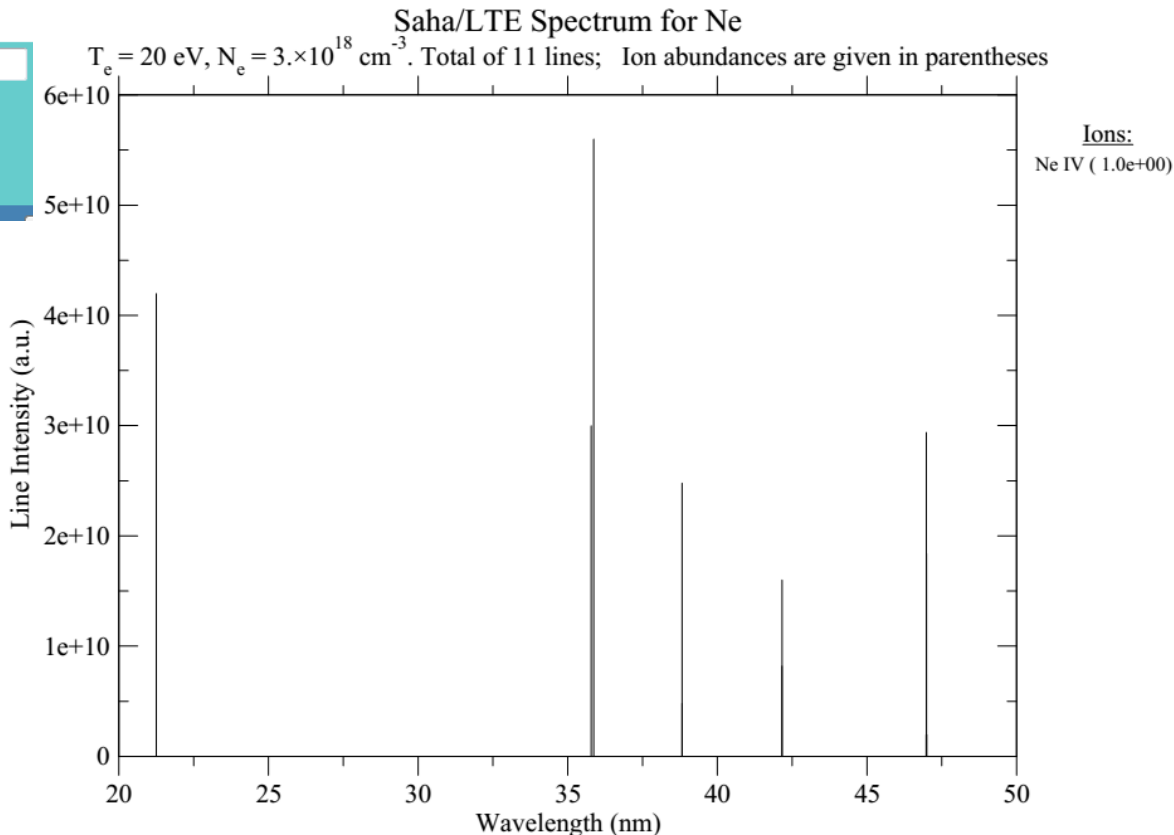
NIST Atomic Spectra Database Lines Form

Spectrum
Ne

Limits for
Wavelengths
Lower: 20
Upper: 50
Wavelength Units: nm

This looks ideal – but there's a problem.

We asked for ALL the ionization states – but specified only strong lines. NIST filters first on the line strength, then includes only those transitions in the calculation. So concludes only Ne +3 in the plasma. Correct, but only if we can get exquisite control of the charge states!



Try to compare Xe +10 and Ne +3...

Plasma	Xe +10 (XI)	Ne +3 (IV)
plasma length	0.01	0.01
radius	0.003	0.003
volume	2.827E-07	2.82743E-07
pressure, mt	100	600
total particles	1.00E+15	6.00E+15
Initial density	3.5E+21	2.1E+22
Pinched radius	2.00E-04	2.00E-04
Volumetric compression	225	225
pinched density	8.0E+23	4.8E+24
N, cgs	8.0E+17	4.8E+18
Joules/pulse	1.4	1.4
Joules/pulse/particle	1.40E-15	2.33E-16
ev/pulse/particle	8753.7	1459.0
Sum of ionizations	806	126
Ratio	10.86	11.58

← Boost the pressure...

Estimate how many eV required to produce the charge state we want by summing ionization energies – compare to Z-pinch eV/particle

← To get a similar ratio...

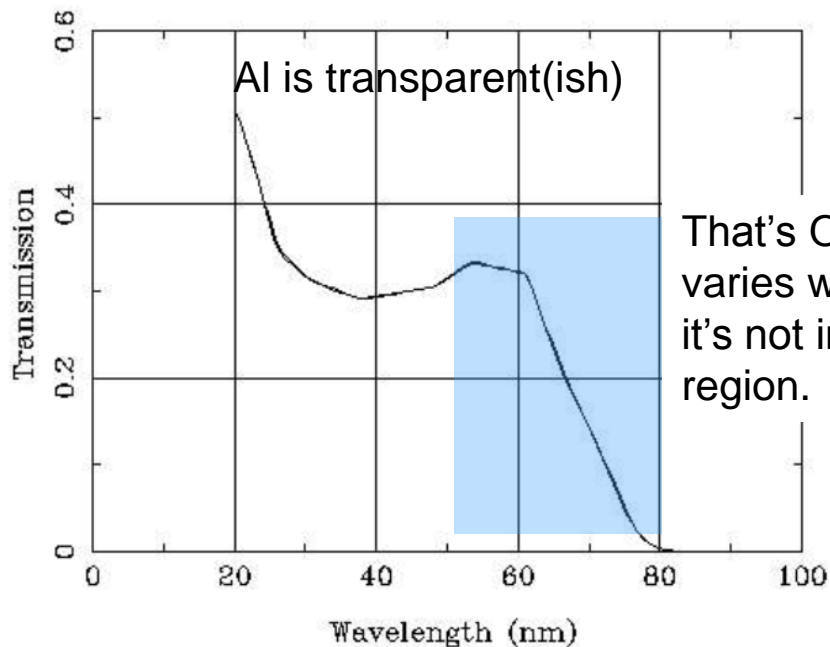
Do the experiment... Run the source in Neon, and measure – how? Need a diagnostic...

- Ideal diagnostic – absolutely calibrated, VUV spectrometer.
Don't have one.
- Simple diagnostic – metal X-ray filters on VAT valves, gas filter, and an X-ray diode. Look at dependence of detected x-ray power as gas filter line density is varied. Compare to Berkeley CXRO data.
- Aluminum has a nice window in this range. Two 500 nm thick filters allows in situ transmission calibration.
- Helium has a convenient cutoff well matched to the Aluminum windows.
- Note – technique will be biased to detecting higher energy lines
 - Lower energy lines will decrease with pressure more rapidly, leaving the higher energies to be detected...

A subtle point.. what about 50-80 nm?

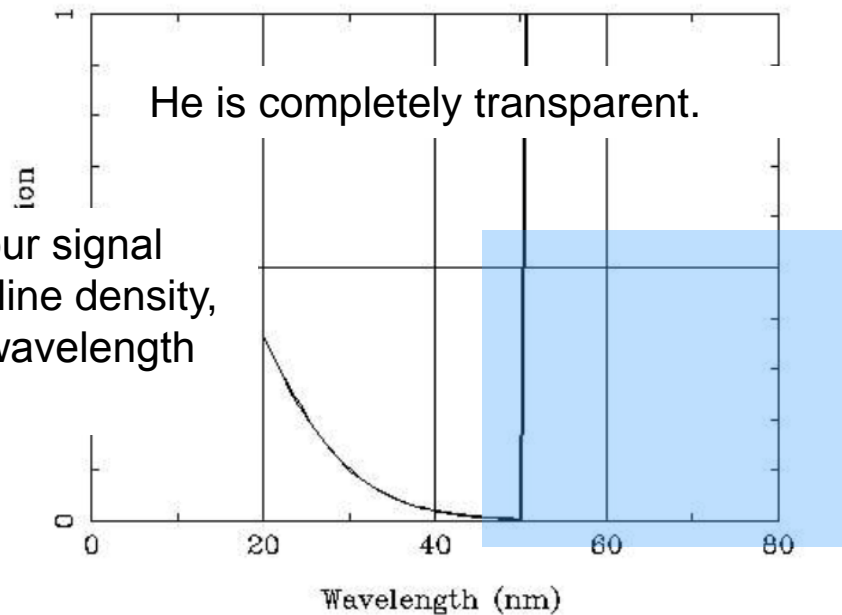
Filter Transmission

Al Density=2.6989 Thickness=0.5 microns

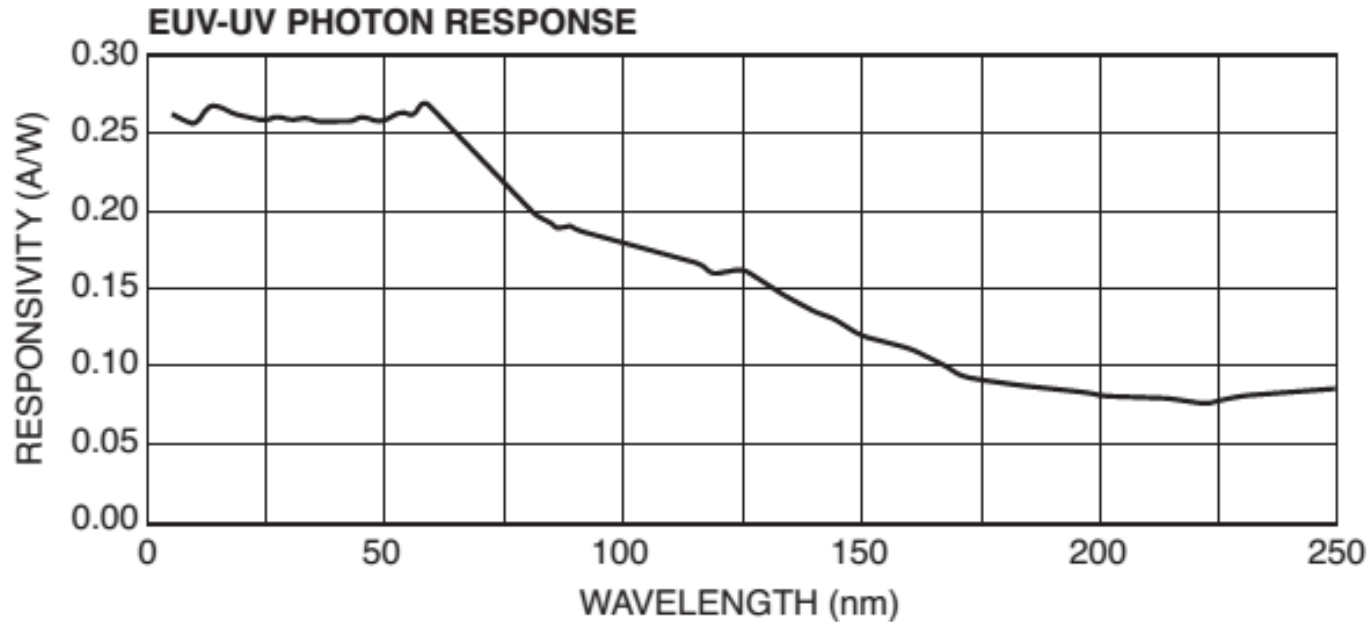


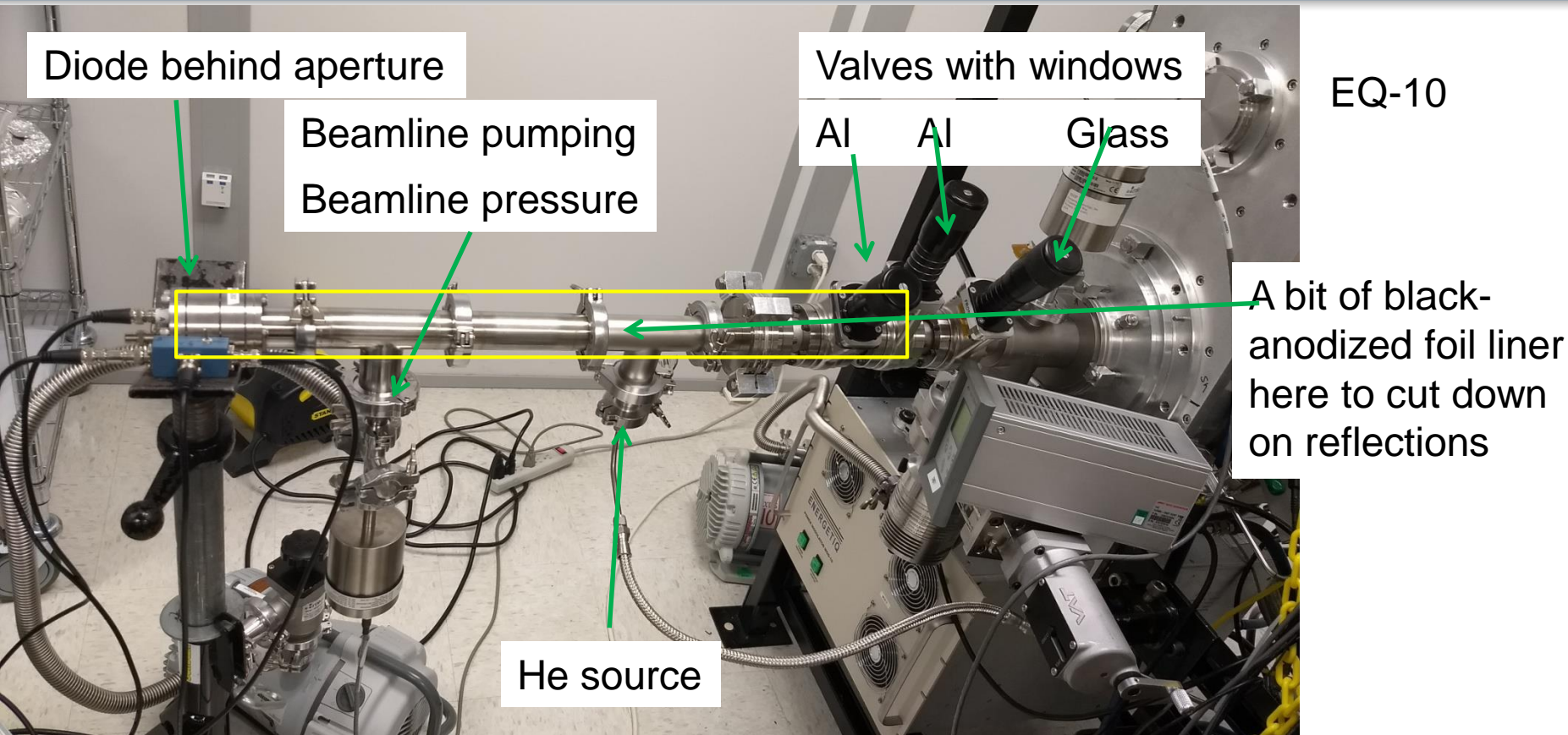
Gas Transmission

He Pressure=1. Path=22.8 cm



Diode reponse (AXUV-100) is quite flat over the region of interest....





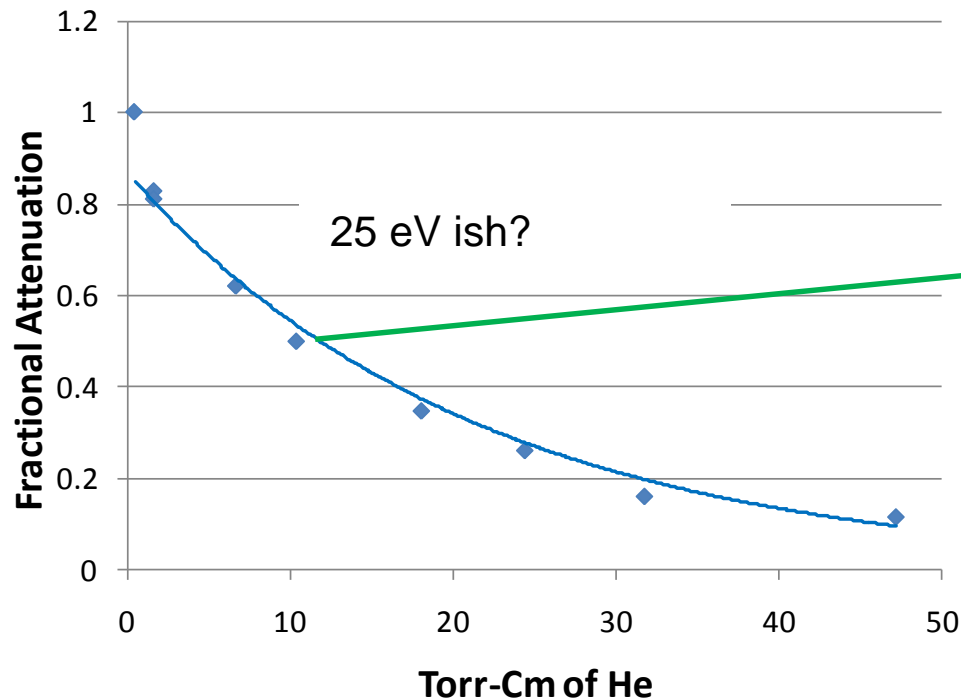
Process --

- Ignite plasma in Argon; switch to Neon...(Neon ionization potential a bit too high?)
- Establish operating condition
- Evacuate beamline
- Measure power with 2-foil method (calibrates foil transmission)
- With one foil in, vary beamline He pressure
- Note diode signal, pressure

Operating point; various constraints -

- Choose a conservative operating point -- about 1.4 J/pulse, 2000 Hz... about 4 kW DC power into pulse shaping system...
- Discover – can't operate at 600 mT pressure. Wrong MFC in the system... max pressure available ~ 170 mT.
 - Could fix this – but it's Tuesday night and my flight is Thursday.... And it's supposed to be a theory presentation anyway...
- Move on. Make the power measurement. Get ~ 19 W with double-foil method.
- Now vary helium pressure in the beamline, note signal decay...

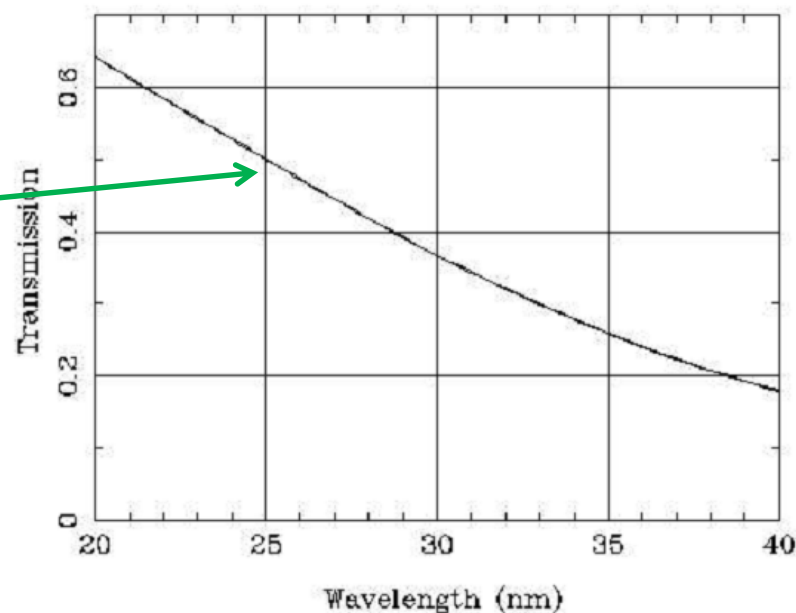
Attenuation due to He in beamline



25 eV ish?

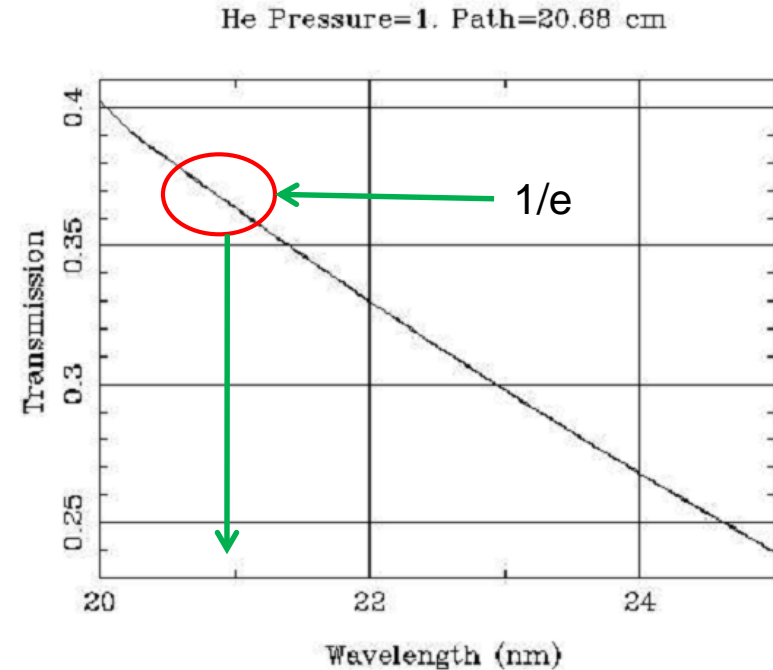
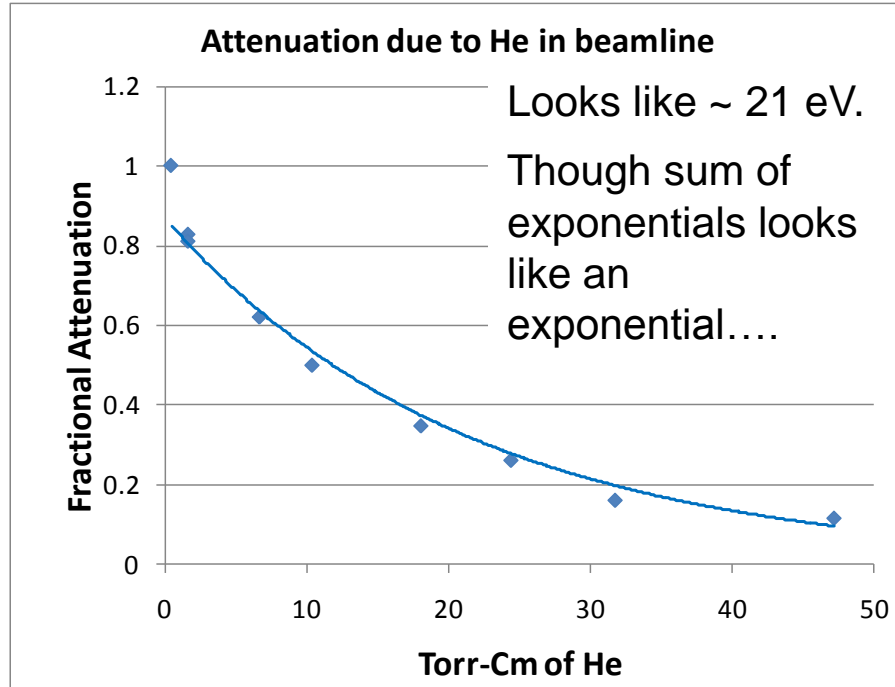
Gas Transmission

He Pressure=1. Path=10. cm



More carefully.. Fit the exponential – get $1/e$ of 20.68 torr-cm, \rightarrow 21 eV radiation

Gas Transmission



We have way too much energy/particle to get Ne IV... what's a more likely charge state?

Plasma	Xe +10 (XI)	Ne +3 (IV)	Ne +3 (IV)	Ne +6 (VII)	Ne +7 (VIII)	Ne+8 (IX)
pressure, mt	100	600	172	172	172	172
ev/pulse/particle	8753.7	1459.0	5089.4	5089.4	5089.4	5089.4
Sum of ionizations	806	126	126	507	714	953
Ratio	10.86	11.58	40.39	10.04	7.13	5.34
	Xenon Reference state	Tried for this. Couldn't raise density enough.		This state probably can't exist -- burned through.		Helium-like Ne -- can't burn past this state - ionization barrier.
		These two states most likely.		Both have lines consistent with measurements		

Calculate with all states; get these ionization fractions..

Ne IV (3.2e-08)

Ne V (3.5e-05)

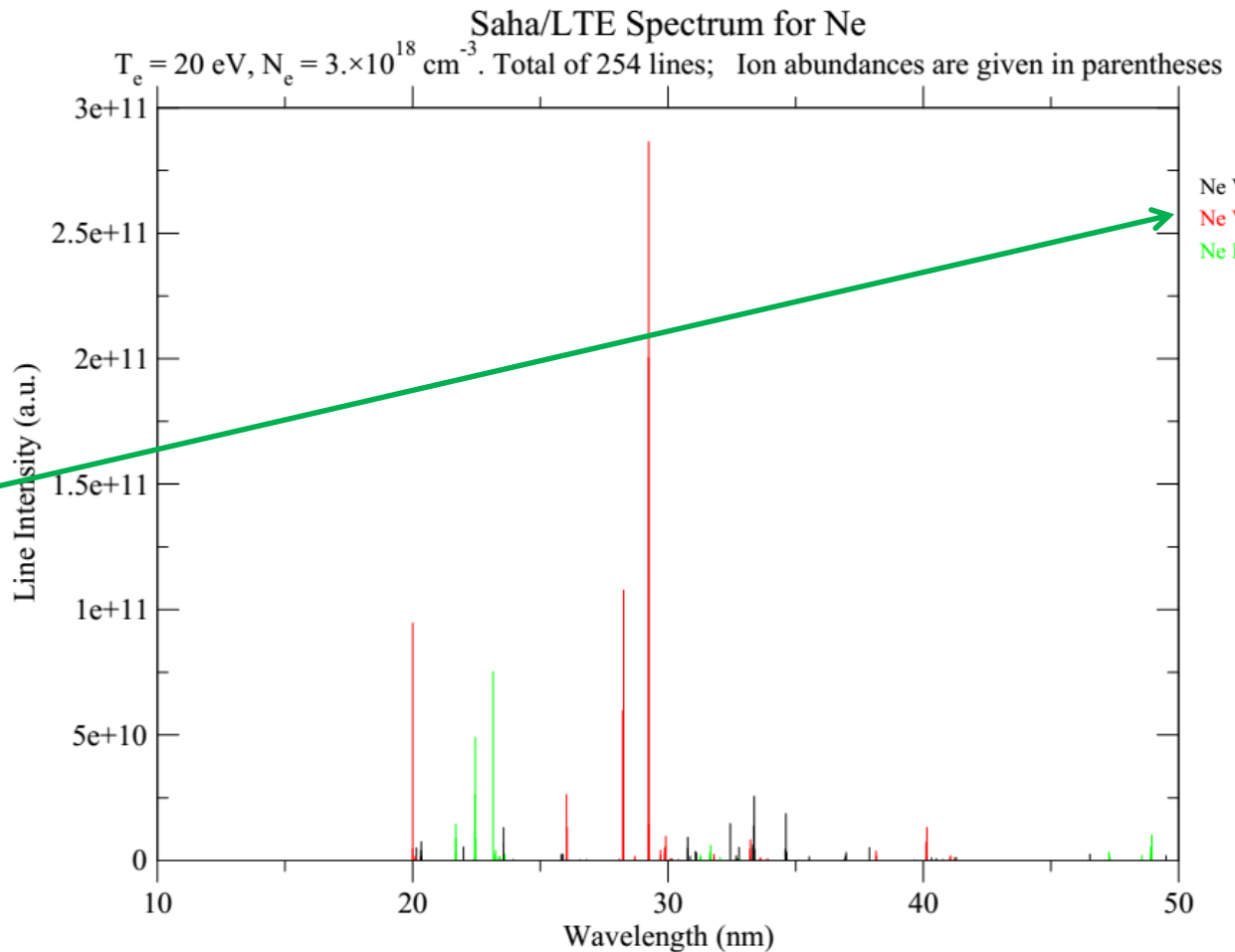
Ne VI (7.2e-03)

Ne VII (2.1e-01)

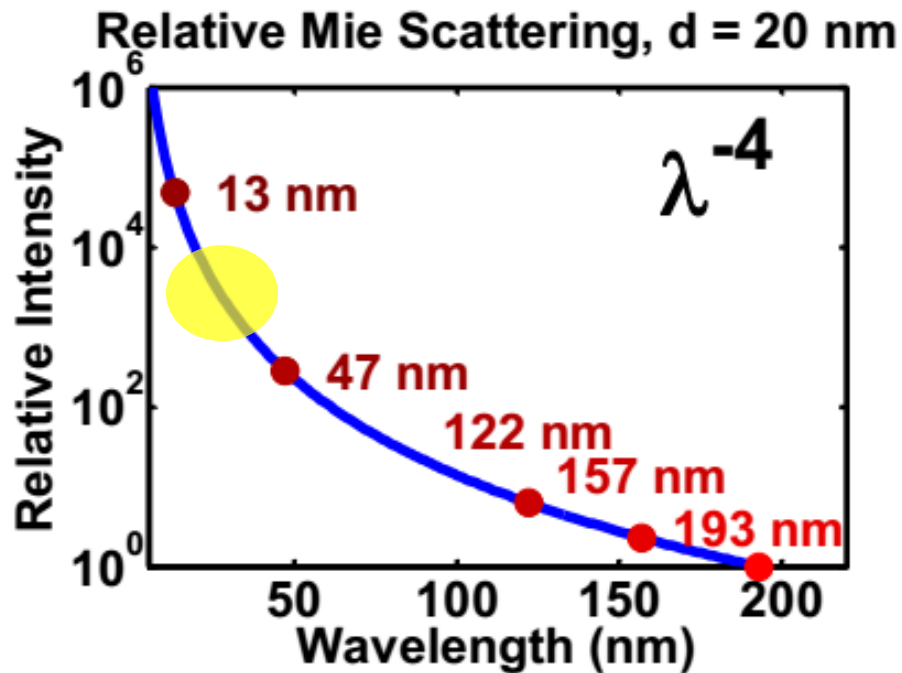
Ne VIII (6.3e-01)

Ne IX (1.5e-01)

Measurement is not inconsistent with a NIST spectral calculation, using all lines and three most likely charge states.



Conclusions - Is this useful for metrology?



Theoretically, we should be able to produce significant power between 20 and 50 nm.

In a very crude and un-optimized experiment, we estimate 19 W/2PI in this range, with 4 kW input power (could go to 7 kW)

If we can increase pressure and avoid wasting energy on high charge states, 100 W should be possible.

Ongoing...

- Replace the MFC, get to higher pressures.
- Need more information on preferred wavelengths – multiple lines? Broadband?
 - (is there a market??)
- Is wavelength selection necessary? Do multilayer mirrors exist that might work?
- Where can we find an appropriate spectrometer?
- I also took data with an Argon plasma. Actually saw more power than with Neon – but NIST is missing most of the rate coefficients, so couldn't connect theory → experiment. Need a spectrometer to proceed.

Thank you for your attention!